

Heterotic string phenomenology: Moduli stabilisation and non-maximally-symmetric spacetime

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Outline

1 Why heterotic string theory?

2 Mathematical background

- Homology
- Differential forms and cohomology

3 Moduli stabilisation

- Moduli stabilisation overview
- Mirror symmetry
- Non-CY manifolds

4 Non-maximally-symmetric spacetime

- Domain wall vacuum
- Extension: cosmic strings and black holes

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Why the heterotic string?

Some reasons to study heterotic string theory:

- String theory: consistent theory of **quantum gravity**!
- Comes equipped with an $E_8 \times E_8$ (or $SO(32)$) gauge group
 \Rightarrow good framework for **grand unified models**.
- Many **candidate Standard Model** compactifications known.
- Calabi–Yau compactification gives $\mathcal{N} = 1$ SUSY in $d = 4$
 \rightarrow suitable for **MSSM-style models**.
- Appealing **mathematical framework**, reasonably well-studied.

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Homology

For a (compact) manifold M :

- A **p -cycle** is a p -dimensional submanifold with no boundary,

$$\partial c_p = 0 .$$

- A **p -boundary** is the boundary of a $(p+1)$ -dimensional submanifold,

$$b_p = \partial d_{p+1} .$$

- **NOTE:** A boundary has no boundary \Rightarrow a p -boundary is a p -cycle.

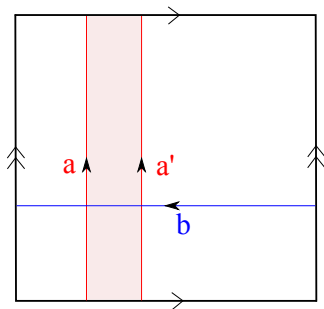
The **p th homology group** is defined as

$$H_p(M) = \frac{\{p\text{-cycles on } M\}}{\{p\text{-boundaries on } M\}} .$$

It classifies the p -cycles that are not boundaries.

Homology example: torus T^2

- a and a' belong to the same homology class (since $a - a'$ is the boundary of the shaded region).
- a and b belong to different homology classes. They cannot be deformed into each other.
- Each 1-cycle can wrap the torus an integer number of times.
- The first homology group, $H_1(T^2) = \mathbb{Z} \oplus \mathbb{Z}$.



Differential forms

- Recall that $\{dx^\mu\}$ form a basis of covariant vectors.
- The **wedge product** is the antisymmetric tensor product, e.g.

$$dx^1 \wedge dx^2 = dx^1 \otimes dx^2 - dx^2 \otimes dx^1 .$$

- A **p-form** is a totally antisymmetric, covariant tensor of rank p :

$$\omega_p = \frac{1}{p!} \omega_{\mu_1 \dots \mu_p} dx^{\mu_1} \wedge \dots \wedge dx^{\mu_p} .$$

- The **exterior derivative** is defined as

$$d\omega_p = \frac{1}{p!} (\partial_\rho \omega_{\mu_1 \dots \mu_p} dx^\rho) \wedge dx^{\mu_1} \wedge \dots \wedge dx^{\mu_p} .$$

Differential forms: example (electromagnetism)

- Define gauge potential 1-form, field strength 2-form,

$$A = A_\nu dx^\nu, \quad F = \frac{1}{2} F_{\mu\nu} dx^\mu \wedge dx^\nu.$$

- The usual relation,

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu,$$

can be written as

$$F = dA.$$

- Similarly, two of Maxwell's equations in tensor form,

$$\partial_\rho F_{\mu\nu} + \partial_\mu F_{\nu\rho} + \partial_\nu F_{\rho\mu} = 0,$$

are now simply

$$dF = 0.$$

Cohomology

- A p -form ω_p is **closed** if

$$d\omega_p = 0 .$$

- A p -form β_p is **exact** if there is a $(p-1)$ -form α_{p-1} such that

$$\beta_p = d\alpha_{p-1} .$$

Electromagnetism example:

- $dF = 0 \Rightarrow F$ is closed;
- $F = dA \Rightarrow F$ is also exact.

On any (compact) manifold M , the **p th cohomology group** is

$$H^p(M) = \frac{\{\text{closed } p\text{-forms on } M\}}{\{\text{exact } p\text{-forms on } M\}} .$$

It classifies the closed p -forms that are not exact.

de Rham's theorem

- Homology and cohomology groups are dual to each other,

$$H_p(M) \cong H^p(M) .$$

- Explicitly this means that

cycles that are not boundaries \Leftrightarrow closed forms that are not exact.

- Or, heuristically,

non-trivial topology \Leftrightarrow non-trivial tensor structure.

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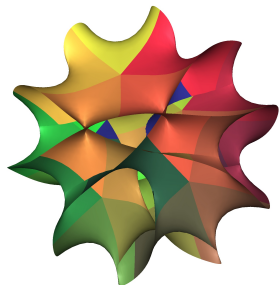
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Calabi–Yau compactification

- Superstring theory is self-consistent only in 10 spacetime dimensions.
- Assume the extra 6 spatial dimensions are **compactified**.
- Lots of supersymmetry in $d = 10$ → want to break most of it.
- Amount of broken SUSY \Rightarrow **holonomy group** of compactification manifold.
- **Holonomy group**: set of all possible rotations of a vector after transport around a closed curve on the manifold.
- Maximum holonomy is $SO(6) \cong SU(4) \Rightarrow$ no SUSY preserved.
- **Calabi–Yau manifold**: $SU(3)$ holonomy \Rightarrow 1/4 SUSY preserved.



What are moduli?

Compactification gives rise to scalar fields called **moduli**.

Three types of moduli:

- The (axio-)dilaton S
→ sets the string coupling, always present in string theory;
- Kähler moduli T^i
→ parametrise closed 2-forms \Rightarrow size of 2-cycles in the geometry;
- Complex structure moduli Z^a
→ closed 3-forms \Rightarrow 3-cycles, “shape” of compactification.

Moduli are flat directions in the potential — this is catastrophic!

Moduli need masses to prevent decompactification, 5th forces, etc.

→ problem of **moduli stabilisation**.

Moduli stabilisation: Type IIB example

- Example: moduli stabilisation in type IIB string theory.
- Theory contains R-R 3-form flux F_3 and NS-NS 3-form flux H_3 .
- R, NS: periodicity conditions on string excitations.
- Compactify such that on the manifold, F_3 and H_3 are non-zero
→ **flux compactification**.
- In the right combination, dilaton and all complex structure moduli can be stabilised (note: 3-forms stabilise 3-cycles).
- Kähler moduli remain unstabilised, can fix with eg.
 - non-perturbative effects (KKLT),
 - non-perturbative effects and perturbative α' corrections (LVS).
- **All moduli stabilised!**

Problems with heterotic moduli stabilisation

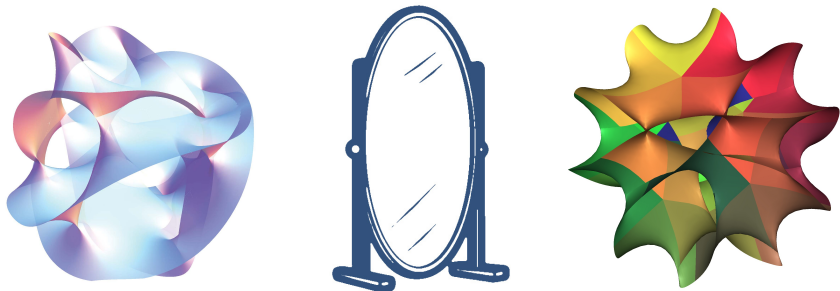
- In heterotic string theory, only have NS-NS flux H_3 .
- Can stabilise complex structure moduli... what then?
- Dilaton can be stabilised by gaugino condensation.
- No other non-perturbative effects, no other options for flux quantisation.
- In fact, problem is even worse:

Strominger, 1986

If a heterotic compactification on a manifold Y has a **maximally symmetric** (i.e. Poincaré) vacuum and non-vanishing H_3 , Y is non-Calabi–Yau.

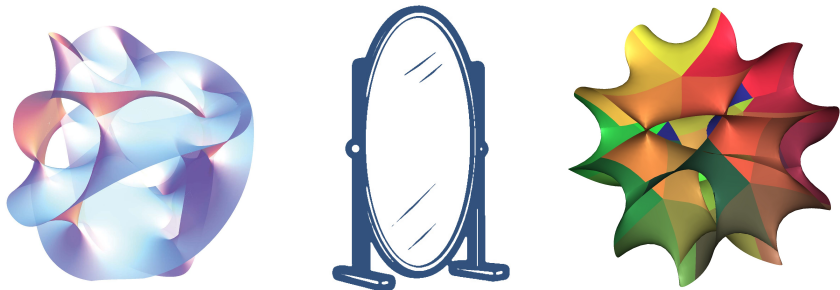
- Hence for a Calabi–Yau compactification, $H_3 = 0$!

Mirror symmetry



- A related issue is **mirror symmetry**.
- For every Calabi–Yau manifold Y , there is a mirror Calabi–Yau \tilde{Y} .
- Mirror symmetry exchanges Kähler moduli T^i and complex structure moduli Z^a .

Mirror symmetry and flux compactifications



- Type IIA compactified on $Y \leftrightarrow$ type IIB compactified on \tilde{Y} .
- Flux compactifications: R-R flux $F_3 \leftrightarrow F_0, F_2, F_4, F_6$.
- Problem: no obvious mirror dual for NS-NS flux H_3 !

What is an $SU(3)$ structure manifold?

- Mirror dual: manifold with $SU(3)$ structure, but not Calabi-Yau
hep-th/0008142 (Vafa), hep-th/0211102 (Gurrieri, Louis, Micu, Waldram).
- **$SU(3)$ structure**: there is a globally-defined spinor η that leaves 1/4 of the SUSY unbroken.
- Calabi-Yau case: η is covariantly constant with respect to the Levi-Civita connection ∇ .
- Non-CY case: $\nabla\eta \sim T^0\eta$ (note: Γ matrices/tensor indices suppressed).
- T^0 is the **intrinsic torsion** of the manifold.
- Torsion on 2-cycles can stabilise Kähler moduli T^i .

A choice

There are two options:

Option 1:

Study the moduli space of Poincaré-invariant compactifications on SU(3)-structure manifolds.

- Has been studied in eg. hep-th/0408121 (Gurrieri, Lukas, Micu), hep-th/0507173 (de Carlos, Gurrieri, Lukas, Micu).
- Torsion quantisation understood from mirror symmetry on so-called **half-flat manifolds**.
- Can stabilise all moduli, but difficult to tune for consistent solutions/GUT gauge couplings. Many axions remain unstabilised.

Option 2:

Heterotic Calabi–Yau flux compactifications, but break maximal symmetry (Poincaré invariance) of $d = 4$ spacetime.

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Domain wall vacuum

- Assume heterotic Calabi–Yau compactification with H -flux.
- Maximal symmetry in $d = 3 + 1$ broken!
- There exist 1/2-BPS domain wall solutions
1305.0594 (Klaput, Lukas, Svanes).
- 1/2-BPS: 2 of the 4 SUSY generators in $d = 4$, $\mathcal{N} = 1$ unbroken.
- $d = (2 + 1)$ Poincaré symmetry preserved; DW breaks symmetry in transverse y direction.
- Moduli satisfy flow equations in the y coordinate.

BUT... that's not realistic!

- However, Kähler moduli have not yet been stabilized explicitly.
- Doing so may “uplift” to a Minkowski vacuum (shown to happen in certain half-flat cases).

Extension

An idea I'm currently working on, in collaboration with Eirik Svanes (LPTHE, Paris) and Cyril Matti (City University, London):

- The 1/2-BPS domain wall breaks Poincaré invariance along one coordinate.
- Mathematically, the $SU(3)$ structure is **fibred** over an “interval”.
- Example of an interval: the real line \mathbb{R} .

Proposition:

Fibre the $SU(3)$ structure over an interval that is not a Cartesian coordinate direction.

Examples:

- Cylindrical (ρ, ϕ, z) , fibre along $\rho \Rightarrow$ **cosmic string**;
- Spherical (r, θ, ϕ) , fibre along $r \Rightarrow$ **black hole**.

Cosmic strings

- Cylindrical polar coordinates (ρ, ϕ, z) : fibre along ρ
 \Rightarrow **cosmic string**.
- Appears to also be 1/2-BPS.
- Flow equations for moduli have the same structure as in the domain wall case.

Issues:

- **That's still not realistic** \Rightarrow again may need to uplift to Minkowski.
- Deeper structure: relation to intersecting domain walls?
- Possible 1/4-BPS solutions?
- Work in progress!

Black holes

- Spherical polar coordinates (r, θ, ϕ) : fibre along r
⇒ “black hole” solution.
- Note: this is a toy model — just a naked singularity at the origin.
- Flow equations have analogous structure to the other 2 cases.
- Major advantage: do not necessarily need to uplift to Minkowski!

However

- Toy model does not yet account for curvature, etc.
- Not yet matched to the full 10d heterotic solution.
- Related to triple-intersecting domain walls? 1/8-BPS not possible as only 4 supercharges. . . consistency?

Summary

- String compactifications generate moduli, which must be stabilised. This can be done using fluxes that wrap the cycles in the geometry.
- For the heterotic string, only H_3 present. One solution is to compactify on $SU(3)$ structure manifolds which are not Calabi–Yau.
- An alternative is to sacrifice Poincaré invariance
→ domain wall solutions have been considered.
- We are currently working on extending this to cosmic strings and black holes.